

## Decay of Dusty Fluid Turbulence Before the Final Period in a Rotating System

Md. Nazrul Islam Mondal

Department of Population Science and Human Resource Development,  
 University of Rajshahi, Rajshahi-6205, Bangladesh

**Abstract:** In this study we are concerned with homogeneous dusty fluid turbulence in a rotating system and have considered correlations between fluctuating quantities, at four points. In this case, three and four point correlation equations are used and the set of equations is made determinate by neglecting the quintuple correlations in comparison to the third and fourth order correlation terms. For convenience, the correlation equations are converted to spectral form by taking their Fourier transforms. Finally, integrating the energy spectrum over all wave numbers, the energy decay law of homogeneous dusty fluid turbulence in a rotating system before the final period is obtained.

**Key words:** Deissler's method, correlation, dusty fluid, energy decay law

### INTRODUCTION

In recent year, the motion of dusty viscous fluids in a rotating system has developed rapidly. The motion of dusty fluid occurs in the movement of dust-laden air, in problems of fluidization, in the use of dust in a gas cooling system and in the sedimentation problem of tidal rivers. When the motion is referred to axes, which rotate steadily with the bulk of the fluid, the coriolis force and centrifugal force must be supposed to act on the fluid. The coriolis force due to rotation plays an important role in a rotating system of turbulent flow while the centrifugal force with the potential is incorporated into the pressure. Deissler<sup>[1,2]</sup> generalized a theory "Decay of homogeneous turbulence for times before the final period". Saffman<sup>[3]</sup> derived an equation that described the motion of a fluid containing small dust particles. Dixit and Upadhyay<sup>[4]</sup>, Kishore and Dixit<sup>[5]</sup>, Kishore and Singh<sup>[6]</sup> discussed the effect of Coriolis force on acceleration covariance in ordinary and MHD turbulent flows. Shimomura and Yoshizawa<sup>[7]</sup>, Shimomura<sup>[8,9]</sup> also discussed the statistical analysis of turbulent viscosity, turbulent scalar flux and turbulent shear flows respectively in a rotating system by two-scale Direct-interaction approach. Kishore and Upadhyay<sup>[10]</sup> studied the decay of MHD turbulence in a rotating system. Sarker and Islam<sup>[11]</sup> also studied the decay of dusty fluid turbulence before the final period in a rotating system using two and three point correlation equations. By analyzing the above theories we have studied the decay of dusty fluid turbulence before the final period in a rotating system using three and four point correlation equations.

### CORRELATION AND SPECTRAL EQUATIONS

The equations of motion of dusty fluid turbulence in a rotating system at the points  $p$ ,  $p'$ ,  $p''$  and  $p'''$  are

$$\frac{\partial u_i}{\partial t} + \frac{\partial(u_i u_m)}{\partial x_m} = -\frac{1}{\rho} \frac{\partial p}{\partial x_i} + \nu \frac{\partial^2 u_i}{\partial x_m \partial x_m} - 2\varepsilon_{mni} \Omega_n u_i + f(u_i - v_i) \quad (1)$$

$$\frac{\partial u'_j}{\partial t} + \frac{\partial(u'_j u'_m)}{\partial x'_m} = -\frac{1}{\rho} \frac{\partial p'}{\partial x'_j} + \nu \frac{\partial^2 u'_j}{\partial x'_m \partial x'_m} - 2\varepsilon_{pmj} \Omega_p u'_j + f(u'_j - v'_j) \quad (2)$$

$$\frac{\partial u''_k}{\partial t} + \frac{\partial(u''_k u''_m)}{\partial x''_m} = -\frac{1}{\rho} \frac{\partial p''}{\partial x''_k} + \nu \frac{\partial^2 u''_k}{\partial x''_m \partial x''_m} - 2\varepsilon_{qmk} \Omega_q u''_k + f(u''_k - v''_k) \quad (3)$$

$$\frac{\partial u'''_l}{\partial t} + \frac{\partial(u'''_l u'''_m)}{\partial x'''_m} = -\frac{1}{\rho} \frac{\partial p'''}{\partial x'''_l} + \nu \frac{\partial^2 u'''_l}{\partial x'''_m \partial x'''_m} - 2\varepsilon_{mli} \Omega_i u'''_l + f(u'''_l - v'''_l) \quad (4)$$

where the repeated subscript in a term indicates a summation. Here  $u_i$ , turbulent velocity components;  $v_i$ , dust particle velocity components;  $\rho$ , fluid density;  $\nu$ , kinematic viscosity;  $\Omega_m$ , constant angular velocity components;  $\varepsilon_{mni}$ , alternating tensor;  $p$ , instantaneous pressure;  $m_s = \frac{4}{3} \pi R_s^3 \rho_s$ , mass of a single spherical dust particle of radius  $R_s$ ;  $\rho_s$ , constant density of the material in dust particles;  $f = \frac{KN}{\rho}$ , dimensions of frequency;  $K$ ,

Stock's drag resistance;  $N$ , constant number density of dust particle.

Multiplying Eq. 1 by  $u'_j u'_k u'_l$ , Eq. 2 by  $u_j u'_k u'_l$ , Eq. 3 by  $u_j u'_k u''_l$  and Eq. 4 by  $u_j u'_k u'''_l$  then adding and taking ensemble average writing in terms of the independent variables  $r$ ,  $r'$  and  $r''$  as

$$\begin{aligned} & \frac{\partial}{\partial t} \langle u_i u_j' u_k'' u_l''' \rangle - \frac{\partial}{\partial r_m} \langle u_i u_m u_j' u_k'' u_l''' \rangle - \frac{\partial}{\partial r_m'} \langle u_i u_m u_j' u_k'' u_l''' \rangle - \\ & \frac{\partial}{\partial r_m''} \langle u_i u_m u_j' u_k'' u_l''' \rangle + \frac{\partial}{\partial r_m} \langle u_i u_j' u_k'' u_l''' \rangle + \frac{\partial}{\partial r_m'} \langle u_i u_j' u_k'' u_l''' \rangle + \frac{\partial}{\partial r_m''} \langle u_i u_j' u_k'' u_l''' \rangle \\ & + \frac{\partial}{\partial r_m} \left( \begin{aligned} & -\frac{\partial}{\partial r_i} \langle p u_j' u_k'' u_l''' \rangle - \frac{\partial}{\partial r_i'} \langle p u_j' u_k'' u_l''' \rangle \\ & -\frac{\partial}{\partial r_i''} \langle p u_j' u_k'' u_l''' \rangle + \frac{\partial}{\partial r_i} \langle u_i p' u_k'' u_l''' \rangle \\ & + \frac{\partial}{\partial r_i'} \langle u_i u_j' p'' u_l''' \rangle + \frac{\partial}{\partial r_i''} \langle u_i u_j' p'' u_l''' \rangle \end{aligned} \right) \\ & + 2\nu \left( \begin{aligned} & \frac{\partial^2 \langle u_i u_j' u_k'' u_l''' \rangle}{\partial r_m \partial r_m} + \frac{\partial^2 \langle u_i u_j' u_k'' u_l''' \rangle}{\partial r_m \partial r_m'} + \frac{\partial^2 \langle u_i u_j' u_k'' u_l''' \rangle}{\partial r_m \partial r_m''} \\ & + \frac{\partial^2 \langle u_i u_j' u_k'' u_l''' \rangle}{\partial r_m' \partial r_m'} + \frac{\partial^2 \langle u_i u_j' u_k'' u_l''' \rangle}{\partial r_m' \partial r_m''} + \frac{\partial^2 \langle u_i u_j' u_k'' u_l''' \rangle}{\partial r_m'' \partial r_m''} \end{aligned} \right) \quad (5) \\ & - 2 \left( \begin{aligned} & \varepsilon_{mni} \Omega_n \langle u_i u_j' u_k'' u_l''' \rangle + \varepsilon_{pmj} \Omega_p \langle u_i u_j' u_k'' u_l''' \rangle \\ & + \varepsilon_{qm k} \Omega_q \langle u_i u_j' u_k'' u_l''' \rangle + \varepsilon_{mnl} \Omega_l \langle u_i u_j' u_k'' u_l''' \rangle \end{aligned} \right) \\ & + f \left( \begin{aligned} & -\langle v_i u_j' u_k'' u_l''' \rangle - \langle u_i v_j' u_k'' u_l''' \rangle - \langle u_i u_j' v_k'' u_l''' \rangle \\ & - \langle u_i u_j' u_k'' v_l''' \rangle + 4 \langle u_i u_j' u_k'' v_l''' \rangle \end{aligned} \right) \end{aligned}$$

where the following transformations were used:

$$\frac{\partial}{\partial x_m'} = \frac{\partial}{\partial r_m}, \quad \frac{\partial}{\partial x_m''} = \frac{\partial}{\partial r_m'}, \quad \frac{\partial}{\partial x_m'''} = \frac{\partial}{\partial r_m''} \quad \text{and} \quad \frac{\partial}{\partial x_m} = -\frac{\partial}{\partial r_m} - \frac{\partial}{\partial r_m'} - \frac{\partial}{\partial r_m''}$$

In order to convert Eq. 5 to spectral form, we define the following nine-dimensional Fourier transforms:

$$\langle u_i u_j'(\underline{r}) u_k''(\underline{r}') u_l'''(\underline{r}'') \rangle = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \gamma_{ijkl}(\underline{k}, \underline{k}', \underline{k}'') \exp[i(\underline{k} \cdot \underline{r} + \underline{k}' \cdot \underline{r}' + \underline{k}'' \cdot \underline{r}'')] d\underline{k} d\underline{k}' d\underline{k}'' \quad (6)$$

$$\left\langle \begin{aligned} & u_i u_m u_j'(\underline{r}) u_k''(\underline{r}') \\ & (\underline{r}') u_l'''(\underline{r}'') \end{aligned} \right\rangle = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \gamma_{imjkl}(\underline{k}, \underline{k}', \underline{k}'') \exp[i(\underline{k} \cdot \underline{r} + \underline{k}' \cdot \underline{r}' + \underline{k}'' \cdot \underline{r}'')] d\underline{k} d\underline{k}' d\underline{k}'' \quad (7)$$

$$\left\langle \begin{aligned} & p u_j'(\underline{r}) u_k''(\underline{r}') \\ & (\underline{r}') u_l'''(\underline{r}'') \end{aligned} \right\rangle = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \delta_{jkl}(\underline{k}, \underline{k}', \underline{k}'') \exp[i(\underline{k} \cdot \underline{r} + \underline{k}' \cdot \underline{r}' + \underline{k}'' \cdot \underline{r}'')] d\underline{k} d\underline{k}' d\underline{k}'' \quad (8)$$

$$\left\langle \begin{aligned} & v_i u_j'(\underline{r}) u_k''(\underline{r}') \\ & (\underline{r}') u_l'''(\underline{r}'') \end{aligned} \right\rangle = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \gamma_{ijkl}(\underline{k}, \underline{k}', \underline{k}'') \exp[i(\underline{k} \cdot \underline{r} + \underline{k}' \cdot \underline{r}' + \underline{k}'' \cdot \underline{r}'')] d\underline{k} d\underline{k}' d\underline{k}'' \quad (9)$$

Similarly,

$$\left\langle \begin{aligned} & u_i v_j'(\underline{r}) u_k''(\underline{r}') \\ & (\underline{r}') u_l'''(\underline{r}'') \end{aligned} \right\rangle = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \gamma_{ijkl}(-\underline{k}, \underline{k}', \underline{k}'') \exp[i(\underline{k} \cdot \underline{r} + \underline{k}' \cdot \underline{r}' + \underline{k}'' \cdot \underline{r}'')] d\underline{k} d\underline{k}' d\underline{k}'' \quad (10)$$

$$\left\langle \begin{aligned} & u_i u_j'(\underline{r}) v_k''(\underline{r}') \\ & (\underline{r}') u_l'''(\underline{r}'') \end{aligned} \right\rangle = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \gamma_{ijkl}(\underline{k}, \underline{k}', \underline{k}'') \exp[i(\underline{k} \cdot \underline{r} + \underline{k}' \cdot \underline{r}' + \underline{k}'' \cdot \underline{r}'')] d\underline{k} d\underline{k}' d\underline{k}'' \quad (11)$$

$$\left\langle \begin{aligned} & u_i u_j'(\underline{r}) u_k''(\underline{r}') \\ & (\underline{r}') v_l'''(\underline{r}'') \end{aligned} \right\rangle = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \gamma_{ijkl}(-\underline{k}, \underline{k}', \underline{k}'') \exp[i(\underline{k} \cdot \underline{r} + \underline{k}' \cdot \underline{r}' + \underline{k}'' \cdot \underline{r}'')] d\underline{k} d\underline{k}' d\underline{k}'' \quad (12)$$

Substituting the preceding relations into Eq. 5, we get

$$\begin{aligned} & \frac{d}{dt} (\gamma_{ijkl}) + 2\nu (\underline{k}^2 + \underline{k}_m \underline{k}_m' + \underline{k}_m \underline{k}_m'' + \underline{k}^2 + \underline{k}_m' \underline{k}_m'' + \underline{k}''^2) \gamma_{ijkl} \\ & = [i(\underline{k}_m + \underline{k}_m' + \underline{k}_m'') \gamma_{imjkl}(\underline{k}, \underline{k}', \underline{k}'') - i \underline{k}_m \gamma_{jmikl}(-\underline{k} - \underline{k}' - \underline{k}'', \underline{k}', \underline{k}'') \\ & - i \underline{k}_m' \gamma_{kmijl}(-\underline{k} - \underline{k}' - \underline{k}'', \underline{k}, \underline{k}'') - i \underline{k}_m'' \gamma_{lmijk}(-\underline{k} - \underline{k}' - \underline{k}'', \underline{k}, \underline{k}'')] \\ & - \frac{1}{\rho} [-i(\underline{k}_i + \underline{k}_i' + \underline{k}_i'') \delta_{jkl}(\underline{k}, \underline{k}', \underline{k}'') + i \underline{k}_i \delta_{ikl}(-\underline{k} - \underline{k}' - \underline{k}'', \underline{k}', \underline{k}'') \\ & + i \underline{k}_i' \delta_{ijl}(-\underline{k} - \underline{k}' - \underline{k}'', \underline{k}, \underline{k}'') + i \underline{k}_i'' \delta_{ijk}(-\underline{k} - \underline{k}' - \underline{k}'', \underline{k}, \underline{k}'')] \quad (13) \\ & - 2(\varepsilon_{nmi} \Omega_n + \varepsilon_{pmj} \Omega_p + \varepsilon_{qm k} \Omega_q + \varepsilon_{mnl} \Omega_l) \gamma_{ijkl} \\ & + f \left[ \begin{aligned} & 4\gamma_{ijkl}(\underline{k}, \underline{k}', \underline{k}'') - \gamma_i \delta_{jkl}(\underline{k}, \underline{k}', \underline{k}'') - \gamma_j \delta_{ikl}(-\underline{k} - \underline{k}' - \underline{k}'', \underline{k}', \underline{k}'') \\ & - \gamma_k \delta_{ijl}(-\underline{k} - \underline{k}' - \underline{k}'', \underline{k}, \underline{k}'') - \gamma_l \delta_{ijk}(-\underline{k} - \underline{k}' - \underline{k}'', \underline{k}, \underline{k}'') \end{aligned} \right] \end{aligned}$$

To obtain a relation between the terms on the right side of Eq. 13 derived from the quadruple correlation terms, pressure terms, rotational terms and the dust particle terms in Eq. 5, take the divergence of the equation of motion and combine with the continuity equation to give

$$\frac{1}{\rho} \frac{\partial^2 p}{\partial x_m \partial x_m} = -\frac{\partial^2 (u_m u_n)}{\partial x_m \partial x_n} \quad (14)$$

Multiplying Eq. 14 by  $u_i u_j' u_k'' u_l'''$ , taking ensemble average and writing the resulting equation in terms of the independent variables  $r$  and  $r'$ , gives

$$\begin{aligned} & \frac{1}{\rho} \left( \begin{aligned} & \frac{\partial^2 \langle p u_j' u_k'' u_l''' \rangle}{\partial r_m \partial r_m} + 2 \frac{\partial^2 \langle p u_j' u_k'' u_l''' \rangle}{\partial r_m \partial r_m'} + 2 \frac{\partial^2 \langle p u_j' u_k'' u_l''' \rangle}{\partial r_m \partial r_m''} \\ & + \frac{\partial^2 \langle p u_j' u_k'' u_l''' \rangle}{\partial r_m' \partial r_m'} + 2 \frac{\partial^2 \langle p u_j' u_k'' u_l''' \rangle}{\partial r_m' \partial r_m''} \end{aligned} \right) \\ & + \frac{\partial^2 \langle p u_j' u_k'' u_l''' \rangle}{\partial r_m'' \partial r_m''} = - \left( \begin{aligned} & \frac{\partial^2 \langle u_m u_n u_j' u_k'' u_l''' \rangle}{\partial r_m \partial r_n} + \frac{\partial^2 \langle u_m u_n u_j' u_k'' u_l''' \rangle}{\partial r_m \partial r_n'} \\ & + \frac{\partial^2 \langle u_m u_n u_j' u_k'' u_l''' \rangle}{\partial r_m \partial r_n''} + \frac{\partial^2 \langle u_m u_n u_j' u_k'' u_l''' \rangle}{\partial r_m' \partial r_n} + \frac{\partial^2 \langle u_m u_n u_j' u_k'' u_l''' \rangle}{\partial r_m' \partial r_n'} \\ & + \frac{\partial^2 \langle u_m u_n u_j' u_k'' u_l''' \rangle}{\partial r_m' \partial r_n''} + \frac{\partial^2 \langle u_m u_n u_j' u_k'' u_l''' \rangle}{\partial r_m'' \partial r_n} + \frac{\partial^2 \langle u_m u_n u_j' u_k'' u_l''' \rangle}{\partial r_m'' \partial r_n'} \\ & + \frac{\partial^2 \langle u_m u_n u_j' u_k'' u_l''' \rangle}{\partial r_m'' \partial r_n''} \end{aligned} \right) \quad (15) \end{aligned}$$

The Fourier transform of Eq. 15 is

$$-\frac{1}{\rho} \delta_{ijkl} = \frac{\left( \begin{aligned} & \underline{k}_m \underline{k}_n + \underline{k}_m \underline{k}_n' + \underline{k}_m \underline{k}_n'' + \underline{k}_m' \underline{k}_n + \underline{k}_m' \underline{k}_n' \\ & + \underline{k}_m' \underline{k}_n'' + \underline{k}_m'' \underline{k}_n + \underline{k}_m'' \underline{k}_n' + \underline{k}_m'' \underline{k}_n'' \end{aligned} \right) \gamma_{mnijl}}{\left( \underline{k}^2 + 2\underline{k}_m \underline{k}_m' + 2\underline{k}_m \underline{k}_m'' + \underline{k}^2 + 2\underline{k}_m' \underline{k}_m'' + \underline{k}''^2 \right)} \quad (16)$$

Equation 13 and 16 are the spectral equations corresponding to the four point correlation equations. The spectral equations corresponding to the three point correlation equations are

$$\begin{aligned} & \frac{d}{dt}(k_k \beta_{iik}) + 2\nu(k^2 + k_1 k_1' + k'^2) k_k \beta_{iik} \\ &= ik_k (k_1 + k_1') \beta_{iik}(k, k') - ik_k k_1 \beta_{iik}(-k - k', k') \quad (17) \\ & - ik_k k_1' \beta_{iik}(-k - k', k) - \frac{1}{\rho} \left[ -ik_k (k_1 + k_1') \alpha_{ik}(k, k') + ik_k k_1 \alpha_{ik} \right] \\ & - 2k_k [\varepsilon_{mii} \Omega_m + \varepsilon_{nii} \Omega_n + \varepsilon_{qik} \Omega_q] \beta_i \beta_k' + Rfk_k \end{aligned}$$

Where, 
$$R\beta_i \beta_k' \beta_k'' = 3 \langle \beta_i \beta_k' \beta_k'' \rangle - \langle \gamma_i \beta_i'(k) \beta_k''(k') \rangle - \langle \gamma_i \beta_i'(-k - k') \beta_k''(k') \rangle - \langle \gamma_k \beta_k'(-k - k') \beta_i''(k) \rangle$$
 (say)

R is an arbitrary constant and

$$-\frac{1}{\rho} \alpha_{ik} = \frac{k_1 k_m + k_1' k_m + k_1 k_m' + k_1' k_m'}{k^2 + 2k_1 k_1' + k'^2} \beta_{imik} \quad (18)$$

Here the spectral tensors are defined by

$$\langle u_i u_j'(r) u_k''(r') \rangle = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \beta_{ijk}(k, k') \exp[i(k \cdot r + k' \cdot r')] dk dk' \quad (19)$$

$$\langle u_i u_j u_k'(r) u_l''(r') \rangle = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \beta_{ijkl}(k, k') \exp[i(k \cdot r + k' \cdot r')] dk dk' \quad (20)$$

$$\langle pu_j'(r) u_k''(r') \rangle = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \alpha_{jk}(k, k') \exp[i(k \cdot r + k' \cdot r')] dk dk' \quad (21)$$

A relation between  $\beta_{ijk}$  and  $\gamma_{ijkl}$  can be obtained by letting  $r'' = 0$  in Eq. 6 and comparing the result with Eq. 20

$$\beta_{ijk}(k, k') = \int_{-\infty}^{\infty} \gamma_{ijkl}(k, k', k'') dk'' \quad (22)$$

The spectral equation corresponding to the two point correlation equation in presence of dusty fluid in a rotating system is

$$\begin{aligned} & \frac{d}{dt} \phi_{i,i} + (2\nu k^2 - Qf + 2\varepsilon_{mki} \Omega_m + 2\varepsilon_{nki} \Omega_n) \phi_{i,i} \\ &= ik_k \phi_{iki}(k) - ik_k \phi_{iki}(-k) \end{aligned} \quad (23)$$

where  $\phi_{i,i}$  and  $\phi_{j,ji}$  are defined by

$$\langle u_i u_j'(r) \rangle = \int_{-\infty}^{\infty} \phi_{ij}(k) \exp(i\mathbf{k} \cdot \mathbf{r}) dk \quad (24)$$

and

$$\langle u_i u_j u_k'(r) \rangle = \int_{-\infty}^{\infty} \phi_{ij}(k) \exp(i\mathbf{k} \cdot \mathbf{r}) dk \quad (25)$$

The relation between  $\phi_{ij}$  and  $\phi_{ijk}$  obtained by letting  $r' = 0$  in Eq. 6 and comparing the result with Eq. 25 is

$$\phi_{ij}(k) = \int_{-\infty}^{\infty} \beta_{ijk}(k, k') dk' \quad (26)$$

### SOLUTION NEGLECTING QUINTUPLE CORRELATIONS

Equation 16 shows that if the terms corresponding to the quintuple correlations are neglected, then the pressure force terms also must be neglected. Thus neglecting first and second terms on the right side of Eq. 13, the equation can be integrated between  $t_i$  and  $t$  to give

$$\begin{aligned} \gamma_{ijkl} &= (\gamma_{ijkl})_1 \exp \left\{ -2\nu(k^2 + k_m k_m' + k_m k_m'' + k'^2 + k'_m k_m'' + k''^2) \right. \\ & \left. + 2(\varepsilon_{nmi} \Omega_n + \varepsilon_{pmj} \Omega_p + \varepsilon_{qmk} \Omega_q + \varepsilon_{rml} \Omega_r) - Sf \right\} (t - t_i) \end{aligned} \quad (27)$$

where

$$\begin{aligned} S\gamma_{ijkl} &= 4\gamma_{ijkl}(k, k', k'') - \gamma_i \delta_{jkl}(k, k', k'') - \gamma_j \delta_{ikl}(-k - k' - k'', k', k'') \\ & - \gamma_k \delta_{ijl}(-k - k' - k'', k, k'') - \gamma_l \delta_{ijk}(-k - k' - k'', k, k') \end{aligned} \quad (28)$$

is an arbitrary constant and  $(\gamma_{ijkl})_1$  is the value of  $(\gamma_{ijkl})_1$  at  $t = t_i$ . The quantity  $(\gamma_{ijkl})_1$  can be considered also as the value of  $(\gamma_{ijkl})_1$  at small values of  $k, k'$  and  $k''$ , at least for times when the quintuple correlations are negligible.

Equation 22 and 27 can be converted to scalar form by contracting the indices  $i$  and  $j$ , as well as  $k$  and  $l$ . Substitution of Eq. 18, 22 and 27 into the three point scalar Eq. 17 results in

$$\begin{aligned} & k_k \beta_{iik} = (k_k \beta_{iik})_0 \exp \left\{ -2\nu(k^2 + k_1 k_1' + k'^2) + \right. \\ & \left. 2(\varepsilon_{mii} \Omega_m + \varepsilon_{nii} \Omega_n + \varepsilon_{qik} \Omega_q) - Rf \right\} (t - t_0) + \frac{\pi^2 [a]}{v} \\ & \left\{ \omega^{-1} \exp \left[ -\omega^2 \left( \frac{3}{4} k^2 + \frac{1}{2} k_1 \right) \right] + \left\{ 2(\varepsilon_{nmi} \Omega_n + \varepsilon_{pmi} \Omega_p + \varepsilon_{qmk} \Omega_q + \varepsilon_{rmi} \Omega_r) - Sf \right\} (t - t_i) \right\} \\ & + 2 \left( \frac{1}{4} k^2 + \frac{1}{2} k_1 k_1' + \frac{1}{4} k'^2 \right)^{\frac{1}{2}} \exp \left[ -\omega^2 (k^2 + k_1 k_1' + k'^2) + \right. \\ & \left. \left\{ 2(\varepsilon_{nmi} \Omega_n + \varepsilon_{pmi} \Omega_p + \varepsilon_{qmk} \Omega_q + \varepsilon_{rmi} \Omega_r) - Sf \right\} (t - t_i) \right] \\ & \times \int_0^{\omega \left( \frac{1}{4} k^2 + \frac{1}{2} k_1 k_1' + \frac{1}{4} k'^2 \right)^{\frac{1}{2}}} \exp(x^2) dx + \frac{\pi^2 [b]}{v} \left\{ -\omega^{-1} \exp \left[ -\omega^2 \left( \frac{3}{4} k^2 + k_1 k_1' + k'^2 \right) \right] \right. \\ & \left. + \left\{ 2(\varepsilon_{nmi} \Omega_n + \varepsilon_{pmi} \Omega_p + \varepsilon_{qmk} \Omega_q + \varepsilon_{rmi} \Omega_r) - Sf \right\} (t - t_i) \right\} \\ & + k \exp \left[ -\omega^2 (k^2 + k_1 k_1' + k'^2) + \left\{ 2(\varepsilon_{nmi} \Omega_n + \varepsilon_{pmi} \Omega_p + \varepsilon_{qmk} \Omega_q + \varepsilon_{rmi} \Omega_r) - \right. \right. \\ & \left. \left. Sf \right\} (t - t_i) \right] \int_0^{\frac{1}{2} \omega k} \exp(x^2) dx + \frac{\pi^2 [c]}{v} \left\{ -\omega^{-1} \exp \left[ -\omega^2 (k^2 + k_1 k_1' + \frac{3}{4} k'^2) \right] \right. \\ & \left. + \left\{ 2(\varepsilon_{nmi} \Omega_n + \varepsilon_{pmi} \Omega_p + \varepsilon_{qmk} \Omega_q + \varepsilon_{rmi} \Omega_r) - Sf \right\} (t - t_i) \right\} + k' \exp \\ & \left[ -\omega^2 (k^2 + k_1 k_1' + k'^2) + \left\{ 2(\varepsilon_{nmi} \Omega_n + \varepsilon_{pmi} \Omega_p + \varepsilon_{qmk} \Omega_q + \varepsilon_{rmi} \Omega_r) - Sf \right\} (t - t_i) \right] \\ & \int_0^{\frac{1}{2} \omega k'} \exp(x^2) dx \end{aligned} \quad (30)$$

where

$$\omega = [2\nu(t-t_1)]^{\frac{1}{2}}$$

In order to simplify the calculations, we shall assume that  $[a]_i = 0$ ; that is, we assume that a function sufficiently general to represent the initial conditions can be obtained by considering only the terms involving  $[b]_i$  and  $[c]_i$ .

The substitution of Eq. 26 and 30 in Eq. 23 and setting  $E = 2\pi k^2 \phi_{i,j}$  results in

$$\frac{dE}{dt} + (2\nu k^2 + 2\epsilon_{mki}\Omega_m + 2\epsilon_{nki}\Omega_n - Qf)E = W \quad (31)$$

where

$$\begin{aligned} W = & k^2 \int_{-\infty}^{\infty} 2\pi i [k_k \beta_{ik}(k, k') - k_k \beta_{ik}(-k, -k')]_0 \exp\{-2\nu(k^2 + \\ & + k_k^2 k'^2 + k^2) + 2(\epsilon_{mki}\Omega_m + \epsilon_{nki}\Omega_n + \epsilon_{qki}\Omega_q) - Rf\}(t-t_0)\} dk' \\ & + k^2 \int_{-\infty}^{\infty} \frac{2\pi^{\frac{1}{2}} i}{\nu} [b(k, k') - b(-k, -k')]_i \{-w^{-1} \exp[-w^2(\frac{3}{4}k^2 + k_k k'_k + k'^2) \\ & + 2(\epsilon_{nmi}\Omega_n + \epsilon_{pmi}\Omega_p + \epsilon_{qmi}\Omega_q + \epsilon_{rmi}\Omega_r) - Sf](t-t_1)] \\ & + k \exp[-w^2(k^2 + k_k k'_k + k'^2) + \{2(\epsilon_{rmi}\Omega_n + \epsilon_{pmi}\Omega_p + \epsilon_{qmi}\Omega_q \\ & + \epsilon_{rmi}\Omega_r) - Sf\}(t-t_1)] \int_0^{\frac{1}{2}wk} \exp(x^2) dx\} dk' + \quad (32) \\ & k^2 \int_{-\infty}^{\infty} \frac{2\pi^{\frac{1}{2}} i}{\nu} [c(k, k') - c(-k, -k')]_i \{-w^{-1} \exp[-w^2(k^2 + k_k k'_k + \frac{3}{4}k'^2) \\ & + \{2(\epsilon_{rmi}\Omega_n + \epsilon_{pmi}\Omega_p + \epsilon_{qmi}\Omega_q + \epsilon_{rmi}\Omega_r) - Sf\}(t-t_1)] \\ & + k' \exp[-w^2(k^2 + k_k k'_k + k'^2) + \{2(\epsilon_{nmi}\Omega_n + \epsilon_{pmi}\Omega_p + \epsilon_{qmi}\Omega_q \\ & + \epsilon_{rmi}\Omega_r) - Sf\}(t-t_1)] \int_0^{\frac{1}{2}wk'} \exp(x^2) dx\} dk' \end{aligned}$$

The quantity  $E$  is the energy spectrum function, which represents contributions from various wave numbers or eddy sizes to the total energy.  $W$  is the energy transfer function, which is responsible for the transfer of energy between wave numbers.

In order to find the solution completely and following Deissler<sup>[2]</sup>, we assume that

$$(2\pi)^2 i [k_k \beta_{ik}(k, k') - k_k \beta_{ik}(-k, -k')]_0 = -\beta_0 (k^4 k'^6 - k^6 k'^4) \quad (33)$$

For the bracketed quantities in Eq. 32, we let

$$\begin{aligned} \frac{4\pi^{\frac{7}{2}}}{\nu} i [b(k, k') - b(-k, -k')]_i &= \frac{4\pi^{\frac{7}{2}}}{\nu} \quad (34) \\ i [c(k, k') - c(-k, -k')]_i &= -2\gamma_i (k^6 k'^8 - k^8 k'^6) \end{aligned}$$

where the bracketed quantities are set equal in order to make the integrands in Eq. 32 antisymmetric with respect to  $k$  and  $k'$ .

By substituting Eq. 33 and 34 in Eq. 32 remembering that  $dk' = -\pi k'^2 d(\cos\theta) dk'$  and  $k_k k'_k = k k' \cos\theta$ , ( $\theta$  is the angle between vectors  $\underline{k}$  and  $\underline{k}'$ ) and carrying out the integration with respect to  $\theta$ , we get

$$\begin{aligned} W = & \int_0^{\infty} \left[ \frac{\beta_0 (k^4 k'^6 - k^6 k'^4) k k'}{2\nu(t-t_0)} \{ \exp\{-2\nu(k^2 + k k' + k'^2) \right. \\ & + 2(\epsilon_{mki}\Omega_m + \epsilon_{nki}\Omega_n + \epsilon_{qki}\Omega_q) - Rf\}(t-t_0)] - \exp\{-2\nu(k^2 - k k' + k'^2) \\ & + 2(\epsilon_{mki}\Omega_m + \epsilon_{nki}\Omega_n + \epsilon_{qki}\Omega_q) - Rf\}(t-t_0)] - \gamma_i \frac{(k^6 k'^8 - k^8 k'^6) k k'}{\nu(t-t_1)} \\ & \times \left( (\omega^{-1} \exp[-\omega^2(\frac{3}{4}k^2 + k k' + k'^2) + \{2(\epsilon_{nmi}\Omega_n + \epsilon_{pmj}\Omega_p \right. \\ & + \epsilon_{qmi}\Omega_q + \epsilon_{rmi}\Omega_r) - Sf\}(t-t_1)] \\ & - \omega^{-1} \exp[-\omega^2(\frac{3}{4}k^2 - k k' + k'^2) + \{2(\epsilon_{rmi}\Omega_n + \epsilon_{pmj}\Omega_p \\ & + \epsilon_{qmi}\Omega_q + \epsilon_{rmi}\Omega_r) - Sf\}(t-t_1)] \\ & + \omega^{-1} \exp[-\omega^2(k^2 + k k' + \frac{3}{4}k'^2) + \{2(\epsilon_{nmi}\Omega_n + \epsilon_{pmj}\Omega_p \\ & + \epsilon_{qmi}\Omega_q + \epsilon_{rmi}\Omega_r) - Sf\}(t-t_1)] \\ & - \omega^{-1} \exp[-\omega^2(k^2 - k k' + \frac{3}{4}k'^2) + \{2(\epsilon_{rmi}\Omega_n + \epsilon_{pmj}\Omega_p \\ & + \epsilon_{qmi}\Omega_q + \epsilon_{rmi}\Omega_r) - Sf\}(t-t_1)] \\ & + [k \exp[-\omega^2(k^2 - k k' + k'^2) + \{2(\epsilon_{rmi}\Omega_n + \epsilon_{pmj}\Omega_p + \\ & \epsilon_{qmi}\Omega_q + \epsilon_{rmi}\Omega_r) - Sf\}(t-t_1)] \\ & - k \exp[-\omega^2(k^2 + k k' + k'^2) + \{2(\epsilon_{nmi}\Omega_n + \epsilon_{pmj}\Omega_p + \epsilon_{qmi}\Omega_q \\ & + \epsilon_{rmi}\Omega_r) - Sf\}(t-t_1)]] \int_0^{\frac{1}{2}ok} \exp(x^2) dx \\ & + [k' \exp[-\omega^2(k^2 - k k' + k'^2) + \{2(\epsilon_{nmi}\Omega_n + \epsilon_{pmj}\Omega_p \\ & + \epsilon_{qmi}\Omega_q + \epsilon_{rmi}\Omega_r) - Sf\}(t-t_1)] \\ & - k' \exp[-\omega^2(k^2 + k k' + k'^2) + \{2(\epsilon_{rmi}\Omega_n + \epsilon_{pmj}\Omega_p + \epsilon_{qmi}\Omega_q \\ & + \epsilon_{rmi}\Omega_r) - Sf\}(t-t_1)]] \int_0^{\frac{1}{2}ok'} \exp(x^2) dx \quad \left. \right) dk' \quad (35) \end{aligned}$$

where  $\omega = [2\nu(t-t_1)]^{\frac{1}{2}}$

The integrand in this equation represents the contribution to the energy transfer at a wave number  $k$ , from a wave number  $k'$ . The integral is the total contribution to  $W$  at  $k$ , from all wave numbers. Carrying out the indicated integration with respect to  $k'$  in Eq. 35, results in

$$W = W_p + W_\gamma \quad (36)$$

where

$$W_p = - \frac{\left(\frac{\pi}{2}\right)^{\frac{1}{2}} \beta_0}{256\nu^{\frac{15}{2}}(t-t_0)^{\frac{15}{2}}} \exp \left[ \left( -\frac{3}{2}\epsilon^2 \right) \left( 105\epsilon^6 + 45\epsilon^8 - \right) \right] + \left[ \begin{aligned} & 2(\epsilon_{mki}\Omega_m + \epsilon_{nki}\Omega_n \\ & + \epsilon_{qki}\Omega_q) - Rf \end{aligned} \right] (t-t_0) \quad (37)$$

and

$$W_\gamma = -\frac{\gamma_i}{v^{10}(t-t_1)^{10}} \left[ \frac{\pi^{\frac{1}{2}}}{16} \exp\left[(-\eta^2) \left( \frac{3}{128}\eta^{16} + \frac{3}{8}\eta^{14} + \frac{21}{64}\eta^{12} - \frac{105}{16}\eta^{10} - \frac{945}{128}\eta^8 \right) \right] + \left\{ 2(\varepsilon_{nmi}\Omega_n + \varepsilon_{pmj}\Omega_p + \varepsilon_{qmk}\Omega_q + \varepsilon_{rml}\Omega_r) - Sf \right\} (t-t_1) \right] + \frac{2\pi^{\frac{1}{2}}}{\sqrt{3}} \exp\left[ \left( -\frac{4}{3}\eta^2 \right) \left( \frac{160}{19683}\eta^{16} + \frac{40}{729}\eta^{14} - \frac{14}{27}\eta^{12} - \frac{455}{162}\eta^{10} - \frac{35}{18}\eta^8 \right) \right] + \left\{ 2(\varepsilon_{nmi}\Omega_n + \varepsilon_{pmj}\Omega_p + \varepsilon_{qmk}\Omega_q + \varepsilon_{rml}\Omega_r) - Sf \right\} (t-t_1) \right] \quad (38)$$

$$- \frac{\left(\frac{\pi}{2}\right)^{\frac{1}{2}}}{16} \exp\left[ \left( -\frac{3}{2}\eta^2 \right) \int_0^{\frac{\eta}{2}} \exp(y^2) dy \left( \frac{3}{64}\eta^{17} + \frac{3}{4}\eta^{15} + \frac{21}{32}\eta^{13} - \frac{105}{8}\eta^{11} - \frac{945}{32}\eta^9 \right) \right] + \left\{ 2(\varepsilon_{nmi}\Omega_n + \varepsilon_{pmj}\Omega_p + \varepsilon_{qmk}\Omega_q + \varepsilon_{rml}\Omega_r) - Sf \right\} (t-t_1) + \frac{\pi^{\frac{1}{2}}}{2} \exp\left[ \left( -\frac{3}{2}\eta^2 \right) (5.386\eta^8 + 9.118\eta^{10} + 3.1017\eta^{12} + 0.1793\eta^{14} - 0.03106\eta^{16} - 0.004942\eta^{18} - 3.615 \times 10^{-4}\eta^{20} - 1.890 \times 10^{-5}\eta^{22} - 7.561 \times 10^{-7}\eta^{24} - 2.447 \times 10^{-8}\eta^{26} - 6.64 \times 10^{-10}\eta^{28} - 1.55 \times 10^{-11}\eta^{30} \dots) \right] + \left\{ 2(\varepsilon_{nmi}\Omega_n + \varepsilon_{pmj}\Omega_p + \varepsilon_{qmk}\Omega_q + \varepsilon_{rml}\Omega_r) - Sf \right\} (t-t_1) \right]$$

where  $\eta = v^{\frac{1}{2}}(t-t_1)^{\frac{1}{2}}k$  and  $\varepsilon = v^{\frac{1}{2}}(t-t_0)^{\frac{1}{2}}k$

The quantity  $W_\beta$  is the contribution to the energy transfer arising from consideration of the three-point correlation equation;  $W_\gamma$  arises from consideration of the four-point equation. Integration of Eq. 36 over all wave numbers shows  $\int_0^\infty Wdk = 0$  that indicating the expression for  $W$  satisfies the conditions of continuity and homogeneity

In order to obtain the energy spectrum function  $E$ , we integrate Eq. 31 with respect to time. This integration results in

$$E = E_\beta + E_\gamma + E_\gamma \quad (39)$$

where

$$E_\beta = \frac{J_0 \varepsilon^4}{3\pi v^2 (t-t_0)^2} \exp[-2\varepsilon^2] + \{2(\varepsilon_{mki}\Omega_m + \varepsilon_{nij}\Omega_n) - Qf\} (t-t_0) \quad (40)$$

$$E_\gamma = \frac{(2\pi)^{\frac{1}{2}} \beta_0}{256v^2 (t-t_0)^2} \exp\left[ \left( -\frac{3}{2}\varepsilon^2 \right) (-15\varepsilon^6 - 12\varepsilon^8 + \frac{7}{3}\varepsilon^{10} + \frac{16}{3}\varepsilon^{12} - \frac{32}{3\sqrt{2}}\varepsilon^{13} \exp\left[ -\frac{\varepsilon^2}{2} \int_0^{\frac{\varepsilon}{2}} \exp(y^2) dy \right] \right] + \{2(\varepsilon_{mli}\Omega_m + \varepsilon_{nij}\Omega_n + \varepsilon_{qkl}\Omega_q) - Rf\} (t-t_0) \quad (41)$$

$$E_\gamma = -\frac{\gamma_i}{v^{10}(t-t_1)^9} \left\{ \frac{\pi^{\frac{1}{2}}}{32} \exp[(-\eta^2)] \left( \frac{189}{64}\eta^8 + \frac{1029}{256}\eta^{10} + \frac{287}{256}\eta^{12} + \frac{95}{512}\eta^{14} + \frac{71}{512}\eta^{16} - \frac{71}{512}\eta^{18} \exp(-\eta^2) [Ei(\eta^2) - 0.5772] \right) + \{2(\varepsilon_{nmi}\Omega_n + \varepsilon_{pmj}\Omega_p + \varepsilon_{qmk}\Omega_q + \varepsilon_{rml}\Omega_r) - Sf\} (t-t_1) \right\} + \left( \frac{\pi}{3} \right)^{\frac{1}{2}} \exp\left[ \left( -\frac{4}{3}\eta^2 \right) \left( \frac{7}{9}\eta^8 + \frac{497}{324}\eta^{10} + \frac{1001}{1458}\eta^{12} + \frac{761}{4374}\eta^{14} + \frac{1963}{19683}\eta^{16} - \frac{3926}{59049}\eta^{18} \exp\left[ -\frac{2}{3}\eta^2 \right] [Ei\left(\frac{2}{3}\eta^2\right) - 0.5772] \right) \right] + \{2(\varepsilon_{nmi}\Omega_n + \varepsilon_{pmj}\Omega_p + \varepsilon_{qmk}\Omega_q + \varepsilon_{rml}\Omega_r) - Sf\} (t-t_1) \} + \frac{\pi^{\frac{1}{2}}}{2} \exp\left[ \left( -\frac{3}{2}\eta^2 \right) (0.2307\eta^{10} + 0.3632\eta^{12} + 0.1502\eta^{14} + 0.04463\eta^{16} - 0.01326\eta^{18} \exp\left[ -\frac{1}{2}\eta^2 \right] [Ei\left(\frac{1}{2}\eta^2\right) - 0.5772] + 2.459 \times 10^{-3}\eta^{18} + 2.935 \times 10^{-4}\eta^{20} + 2.846 \times 10^{-5}\eta^{22} + 2.52 \times 10^{-6}\eta^{24} + 1.69 \times 10^{-7}\eta^{26} + 1.25 \times 10^{-8}\eta^{28} + 5.80) \times 10^{-10}\eta^{30} + 4.00 \times 10^{-11}\eta^{32} + \dots) \right] + \{2(\varepsilon_{nmi}\Omega_n + \varepsilon_{pmj}\Omega_p + \varepsilon_{qmk}\Omega_q + \varepsilon_{rml}\Omega_r) - Sf\} (t-t_1) \} + \frac{1}{2} \pi^{\frac{1}{2}} \exp\left[ \left( -\frac{3}{2}\eta^2 \right) (1.077\eta^8 + 2.414\eta^{10} + 1.408\eta^{12} + 0.4416\eta^{14} + 0.1898\eta^{16} - 0.0899\eta^{18} \exp\left[ -\frac{1}{2}\eta^2 \right] [Ei\left(\frac{1}{2}\eta^2\right) - 0.5772] + 6.575 \times 10^{-4}\eta^{18} + 3.271 \times 10^{-5}\eta^{20} + 1.270 \times 10^{-6}\eta^{22} + 4.03 \times 10^{-8}\eta^{24} + 1.08 \times 10^{-9}\eta^{26} + 2.50 \times 10^{-11}\eta^{28} + 5.09 \times 10^{-13}\eta^{30} + \dots) \right] + \{2(\varepsilon_{nmi}\Omega_n + \varepsilon_{pmj}\Omega_p + \varepsilon_{qmk}\Omega_q + \varepsilon_{rml}\Omega_r) - Sf\} (t-t_1) \} \quad (42)$$

The quantity  $E_\gamma$  is the energy spectrum function for the final period, where  $E_\beta$  and  $E_\gamma$  are the contributions to the energy spectrum arising from consideration of the three and four point correlation equations, respectively.

Equation 39 can be integrated over all wave numbers to give the total turbulent energy

$$\frac{1}{2} \langle u_i u_i \rangle = \int_0^\infty E dk \quad (43)$$

The result carrying out the integration is, in dimensionless form

$$\frac{\langle u_i u_i \rangle}{2} = \frac{j_0^{\frac{14}{9}} v^{\frac{5}{9}}}{\beta_0^{\frac{5}{9}}} \left[ \frac{1}{32(2\pi)^2} T^{-\frac{5}{2}} \exp\{[Qf - 2(\varepsilon_{mki}\Omega_m + \varepsilon_{nij}\Omega_n)](t-t_0)\} + 0.2296 T^{-7} \exp\{[Rf - 2(\varepsilon_{mli}\Omega_m + \varepsilon_{nij}\Omega_n + \varepsilon_{qkl}\Omega_q)](t-t_0)\} + 6.18 \frac{\gamma_i v^{\frac{5}{9}} j_0^{\frac{5}{9}}}{\beta_0^{\frac{5}{9}}} \left( \frac{t-t_1}{t-t_0} \right)^{\frac{19}{2}} T^{-\frac{19}{2}} \exp\{[Sf - 2(\varepsilon_{nmi}\Omega_n + \varepsilon_{pmj}\Omega_p + \varepsilon_{qmk}\Omega_q + \varepsilon_{rml}\Omega_r)](t-t_1)\} \right] \quad (44)$$

$$\begin{aligned} \langle u^2 \rangle = & AT^{-\frac{5}{2}} \exp[\{Qf - 2(\epsilon_{mkt}\Omega_m + \epsilon_{njk}\Omega_n)\}(t-t_0)] \\ & + BT^{-7} \exp[\{Rf - 2(\epsilon_{mkt}\Omega_m + \epsilon_{njk}\Omega_n + \epsilon_{qk}\Omega_q)\}(t-t_0)] \\ & + CT^{-\frac{19}{2}} \left(\frac{t-t_1}{t-t_0}\right)^{\frac{19}{2}} \exp[\{Sf - 2(\epsilon_{rmi}\Omega_r + \epsilon_{pmj}\Omega_p \\ & + \epsilon_{qmk}\Omega_q + \epsilon_{rmi}\Omega_r)\}(t-t_1)] \end{aligned} \quad (45)$$

where  $\frac{t-t_1}{t-t_0} = 1 - \left(\frac{\gamma_1 v_0^9 J_0^5}{\beta_0^9}\right)^{1/9} \left[\frac{(t_1-t_0)v_0^{81} J_0^{81}}{\beta_0^{81} \gamma_1^9}\right] \frac{1}{T}$

and  $T = \frac{v_0^{\frac{11}{9}} J_0^{\frac{2}{9}} (t-t_0)}{\beta_0^{\frac{2}{9}}}$

and A, B, C are arbitrary constants.

**CONCLUDING REMARKS**

In Equation 44 we obtain the decay law of dusty fluid turbulence in a rotating system before the final period considering three and four point correlation equations after neglecting quintuple correlation terms. The Eq. 44 shows that turbulent energy decays more rapidly in an exponential manner than the energy decay for non-rotating clean fluid. This decay law contains a term,

$T^{-\frac{19}{2}}$  as well as the terms  $T^{-\frac{5}{2}}$  and  $T^{-7}$  along with exponential terms that also contains rotational terms in presence of dust particles. Thus the terms associated with the higher order correlations die out faster than those associated with the lower order ones. The factor

$$\frac{(t-t_1)}{(t-t_0)}$$

occurring in the last term in Eq. 44 will cause that term to decay even faster, so long as  $t_1-t_0 > 0$ .

If the system is non-rotating, we put  $\Omega's = 0$ , the Eq. 45 becomes

$$\begin{aligned} \langle u^2 \rangle = & AT^{-\frac{5}{2}} \exp\{Qf(t-t_0)\} + BT^{-7} \exp\{Rf \\ & (t-t_0)\} + C \left(\frac{t-t_1}{t-t_0}\right)^{\frac{19}{2}} T^{-\frac{19}{2}} \exp\{Sf(t-t_1)\} \end{aligned} \quad (46)$$

Again if the fluid is clean, we put  $f = 0$ , then Eq. 46 becomes

$$\langle u^2 \rangle = AT^{-\frac{5}{2}} + BT^{-7} + C \left(\frac{t-t_1}{t-t_0}\right)^{\frac{19}{2}} T^{-\frac{19}{2}}$$

which is obtained earlier by Deissler<sup>[2]</sup>.

If the higher order correlations were considered in the analysis, it appears that more terms in higher power of T would be added to Eq. 45.

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